Polaris 3-D CdZnTe (CZT) Gamma-Ray Imaging Spectrometers

Zhong He



ISOE-ALARA Symposium, January 12, 2015

Acknowledgements: **DOD**, **DOE** and **DHS**

John M. Palms Outstanding Innovation Began in early 1960s

PROCEEDINGS OF THE INTERNATIONAL SYMPOSIUM ON CADMIUM TELLURIDE A MATERIAL FOR GAMMA-RAY DETECTORS

JUNE 29.30, 1971

STRASBOURG (FRANCE) XXIII-1

THE EVALUATION OF CdTe FOR GAMMA AND X-RAY SPECTROMETERS AND COUNTERS*

J. M. Palms Physics Department Emory University Atlanta, Georgia 30322

Summary

Radiation detectors have been fabricated from high resistivity (semi-insulating) and low resistivity n-type CdTe, vapor grown and grown from the melt. These devices were Al and Au surface barrier detectors having thicknesses from 0.8mm to 1.5mm, and areas from 2mm² to 25mm². The pulse-height response of these devices to low energy gammas and x-rays were studied as a function of detector bias (0 to 1000 volts) and temperature (-75°C to 52°C). The pulse-height spectra of all devices was affected by the presence of trapping and detrapping of charge carriers. All devices showed the mechanisms of trapping and detrapping of charge carriers. The total counting efficiency for most devices was, however, high (>85%). For a typical detector, an All surface barrier detector fabricated from n-type CdTe, the total system resolution was approximately 10.7 keV (FWHM) for the 122 keV gamma from 57 Co. For this device the 136 keV and 122 keV gammas from 57 Co were clearly resolved; the peak-to-valley ratio for the 122 keV line was 4 to 7 and the K_a x-ray escape peaks were clearly visible. The response of other CdTe detectors to gammas and low energy are presented.

Thank you!

Editors : P. SIFFERT, A. CORNET

20 Ph.Ds Graduated



H=D Inc.













Current Orion group members



Zhong He



Yuefeng Zhu



Jim Berry









Josh Mann

Steven Brown Will Koehler Michael Streicher Sean O'Neal









Jiyang Chu David Goodman Bennett Williams Jiawei Xia

Why CdZnTe?

(1) Superior energy resolution



(2) Higher sensitivity per unit volume (higher Z and density) – compactness CdZnTe(48-30-52, 6.0), High-purity Ge(32, 5.32)

(3) Room-temperature operation (no cryogenic cooling) \rightarrow Wide band-gap

Technical challenges

- (1) Severe hole trapping & electron trapping cause signal deficit
- (2) Crystal yield (cost) and non-uniformity

Before 1994



Signal amplitude = gain×($n \cdot e_0$)×(normalized electron drift length z) \propto Energy × z

Improvement Using 3-D Readout Technology



All-events (no selection)



Polaris Systems

Eighteen 2×2×1.5 cm³ CdZnTe detectors (108 cm³, 648 grams = 1.43 lb)

662 keV

0 keV





<u>Performance Goals</u> $\Delta E/E ≤ 1\%$ FWHM (at 662 keV) <u>Real-time</u> γ Imaging + isotope I.D.



Number of photons: 2033

Polaris 1.1 (GMI ASIC) – August 2010



Comparing to Other γ **Spectrometers**



High resolution → High isotope selectivity/ID & low background

Gamma Imaging Capability



Energy Range of Compton Imaging



Polaris-H performs Compton imaging for γ -rays with $E \ge 250 \text{ keV}$

Coded Aperture Imaging at $E \le 250 \text{ keV}$



<u>**Principle</u>**: Recognize unique mask shadows from different incident gamma-ray angles</u>

Real-Time Combined Coded Aperture and Compton Imaging



Advanced Capabilities

Example 1 Directional spectroscopy



Advanced Capabilities

Example 2

Detect and Characterize moving sources

0 Polar Angle (Degrees) 45.0 90.0 135 180 0 Counts = 210Iterations = 0 Mean = 15.736 Polar Angle (Degrees) 45.0 Stdev = 7.53 59.19 90.0 45.77 32.36 135 18.94 5.522 180 0 Counts = 209 Iterations = 0Mean = 15.678 Polar Angle (Degrees) 45.0 Stdev = 7.5158.55 90.0 45.22 - 31.89 135 - 18.56 180 - 5.233 315 270 . 90.0 0 360 225 180 135 45.0

Counts vs Polar Angle (Degrees) vs Azimuthal Angle (Degrees)

Tracking Moving Targets

Azimuthal Angle (Degrees)

Advanced Capabilities

Example 3 3-Dimensional Imaging

3D Imaging with a Moving Detector Experiment Lab Bench `Х У Desk Wall Lab Bench * Desk **#** Cs-137 (30 µCi) Na-22 (30 µCi) Detector Array How to locate sources in 3D?



Azimuthal Angle (Degrees)



X (cm)

Azimuthal Angle (Degrees)

26

Y (cm)

Advanced Capabilities

<u>Example 4</u>

Detection of Source Shielding

Detection of shielded source



Today and Tomorrow



Univ. of Michigan Polaris Technology Licensed to H3D, Inc.











Fiestaware

Compiled by Dr. Chris Wahl of H3D Inc.





6 hours (367 counts, imaging 766 and 1000 keV lines)



12 hours (786 counts, imaging 766 and 1000 keV lines)



Shape of fiestaware start to appear – what you can see from an overnight measurement

Residual Heat Remover (RHR) Pump Room



10.2 minutes 2.3 x 10⁶ counts ~0.4 mrem/hr



RHR Pump Room – ⁵⁸Co



Area source found on pump.

2.1 x 10⁴ counts in ROI (23% of those imageable)



RHR Pump Room – ⁶⁰Co



⁶⁰Co and ⁵⁸Co are in different regions of pump.

3.7 x 10³ counts in ROI (34% of those imageable)



RHR Pump Room – ¹³⁷Cs



See contamination on floor from prior flood. Nearer areas appear hotter. 4.0 x 10⁴ counts in ROI (29% of those imageable)



RHR Pump Room – ¹³⁷Cs



Shielding Verification



Locating Isotopes in Shipping Containers



Co-60; 17 min.

Applications

- National security, homeland security and international nuclear non-proliferation
- Nuclear power (dose reduction, clean-up and inspection)
- Medical imaging (proton cancer therapy)
- Planetary sciences & astrophysics (NASA)
- Safeguard (IAEA)
- Fundamental Physics
- Environmental monitoring

How does an optical camera work?

Object



spectrometer

Principle of a <u>low-energy</u> gamma camera (< 250 keV)



Note: Identity of isotopes are determined by γ -ray spectroscopy

Principle of <u>Polaris</u> technology for higher energy (> 250 keV, ¹³⁷Cs, ⁵⁸Co, ⁶⁰Co) γ-rays

Object



Advanced Algorithm Fil

image**position_sensional** image**position_sensing**io spectrometers

spectrome





